

$e^+e^-$  Physics Above 500 GeV -  
Quantifying the Advantage of  
High Energy

Jan 29, 2002

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Upcoming LC technology discussions will include energy expandability beyond  $\sqrt{s} = 0.5$  TeV

⇒ Physics gains from higher energy running should be fully explored and understood.

To that end it would seem appropriate for this group to undertake a study with the following two objectives:

1. For several physics scenarios, perform a detailed comparison of the physics programs at  $\sqrt{s} = 0.8, 1.0, 1.2,$  and  $1.5$  TeV.
2. Determine if there is a natural energy scale above  $\sqrt{s} = 0.5$  TeV which the next LC should be capable of attaining through energy expansion.

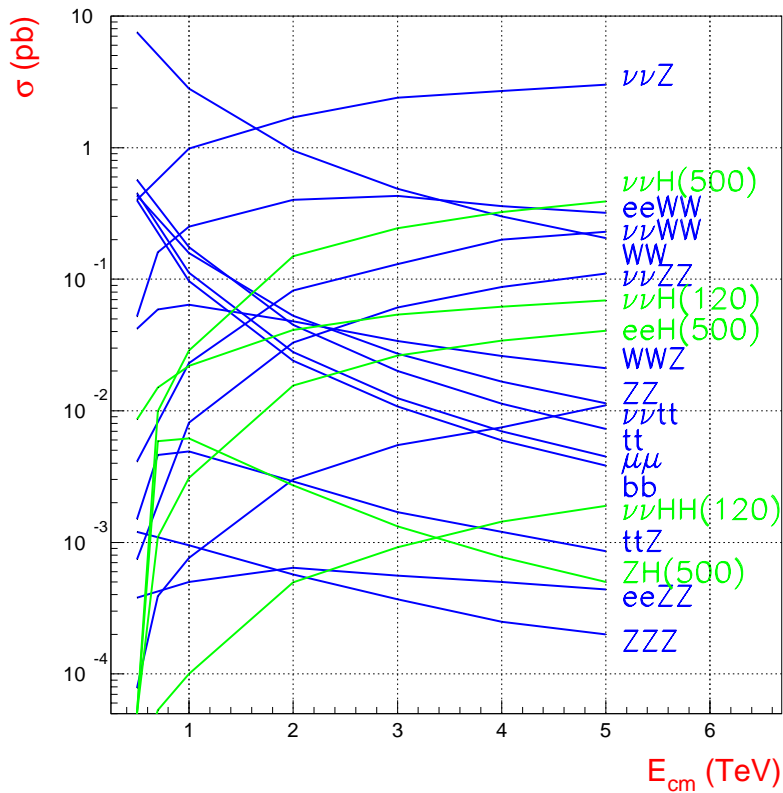
## Physics Scenarios:

- a. SM w/ 120 GeV Higgs Boson
- b. SM w/o Higgs Boson
- c. SUSY Models X,Y,Z

Suggestion is to ignore  $Z'$  resonance production, extra dimensions and other possibilities for now in order to avoid proliferation of physics scenarios.

*Limits* on  $Z'$  bosons, fermion compositeness, etc. can always be included as these are independent of physics scenarios.

# Cross Sections

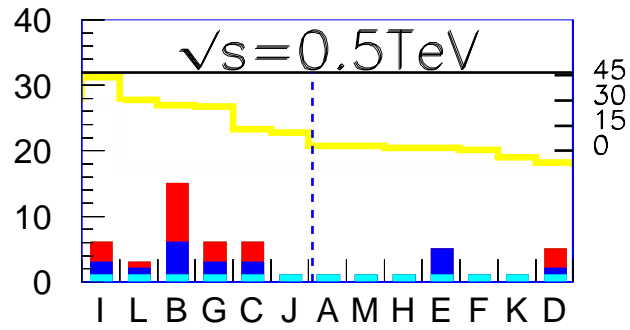
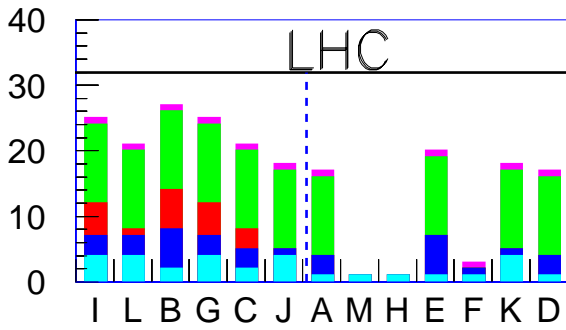


Event Rates/Year ( $1000 \text{ fb}^{-1}$ )	3 TeV $10^3$ events	5 TeV $10^3$ events
$e^+e^- \rightarrow t\bar{t}$	20	7.3
$e^+e^- \rightarrow b\bar{b}$	11	3.8
$e^+e^- \rightarrow ZZ$	27	11
$e^+e^- \rightarrow WW$	490	205
$e^+e^- \rightarrow hZ/h\nu\nu$ (120 GeV)	1.4/530	0.5/690
$e^+e^- \rightarrow H^+H^-$ (1 TeV)	1.5	0.95
$e^+e^- \rightarrow \tilde{\mu}^+\tilde{\mu}^-$ (1 TeV)	1.3	1.0

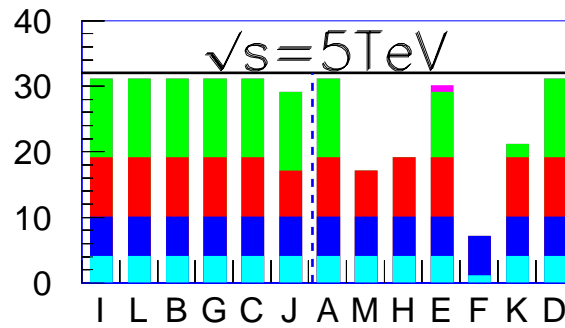
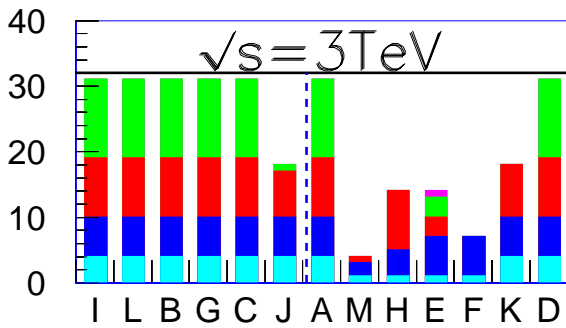
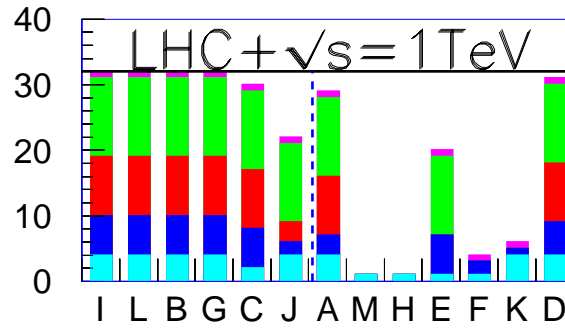
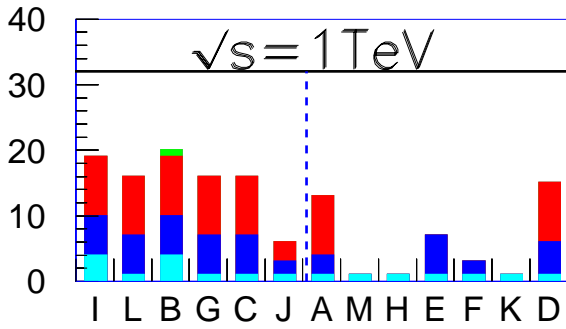
# CMSSM Benchmarks

█ gluino   
 █ squarks   
 █ sleptons   
 █  $\chi^{0,\pm}$    
 █ H

Nb. of Observable Particles

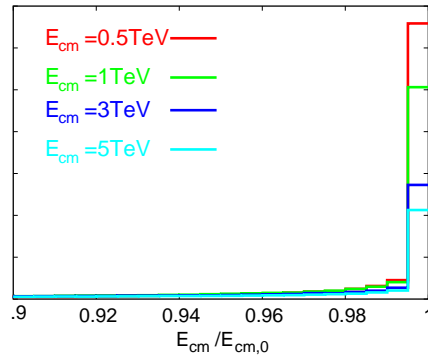


$\delta a_\mu (10^{-10})$



# Luminosity Spectra

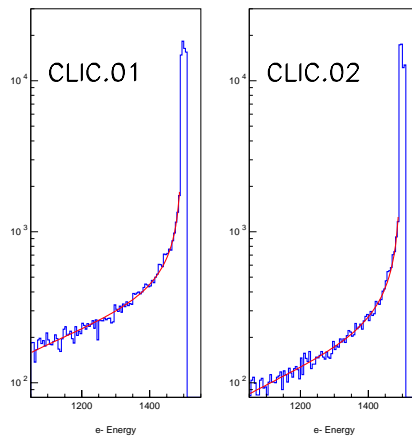
Energy loss due to beam-beam interactions



Luminosity within 1% & 5% of c.m. energy

Energy (TeV)	0.5	1	3	5
$\mathcal{L}$ in 1% $\sqrt{s}$	71%	56 %	30%	25%
$\mathcal{L}$ in 5% $\sqrt{s}$	87%	71 %	42%	34%

Spectra for CLIC studies (sharper  $\leftrightarrow$  high lumi)



CLIC.01:  $\mathcal{L} = 1.05 \times 10^{35}$

CLIC.02:  $\mathcal{L} = 0.40 \times 10^{35}$

## Experimenting at CLIC

Machine parameters lead to challenges for experimenting at CLIC

Beam-Beam effects lead to considerable backgrounds and distortion of the luminosity spectrum (beam strahlung)

FOR 3 TEV AND  $\mathcal{L} = 10^{35} \text{cm}^{-2} \text{s}^{-1}$  (D. Schulte)

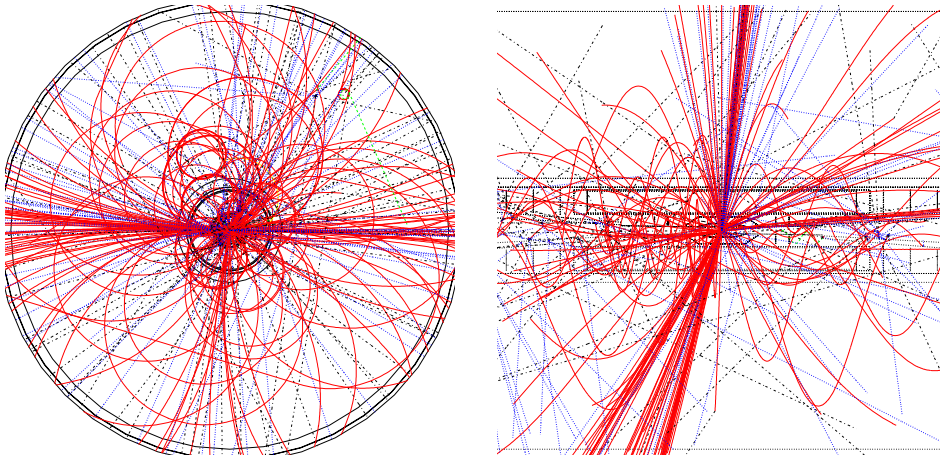
Luminosity/bunch	$10^{-2} \text{nb}^{-1}$
Beam energy spread	1%(FWHM)
Time between bunches	0.67 ns
Average energy loss photons/beam particle	31% 2.3
Number/energy incoh. pairs	$4.6 \cdot 10^5 / 3.9 \cdot 10^4 \text{ TeV}$
Number/energy coh. pairs	$1.4 \cdot 10^9 / 4.4 \cdot 10^8 \text{ TeV}$
Hadronic ( $\gamma\gamma$ ) events, $W_{\gamma\gamma} > 5 \text{ GeV}$	4

- ◆ Coherent pairs disappear in beampipe (backscattering!)
- ◆ Incoherent pairs: suppressed by strong magnetic field of detector

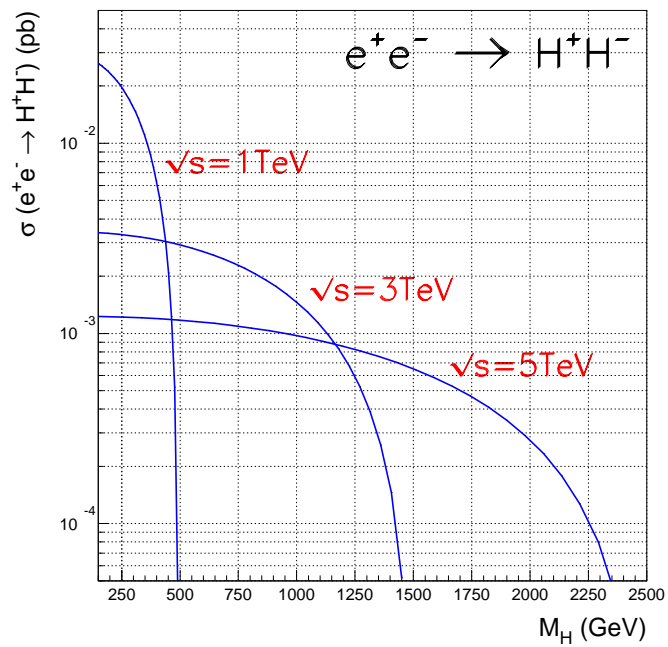
Further backgrounds to consider: neutrons, muons, synchrotron radiation...

$$\sqrt{s} = 3 \text{ TeV}$$

$$e^+e^- \rightarrow H^+H^- \quad M_H = 900 \text{ GeV}$$



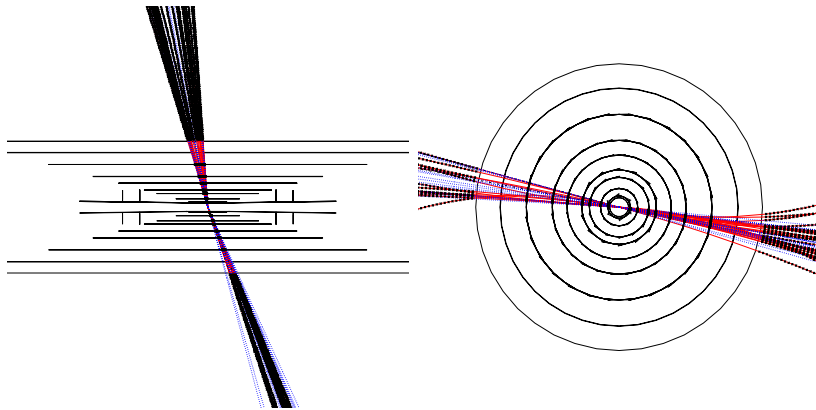
Cross section as function of mass



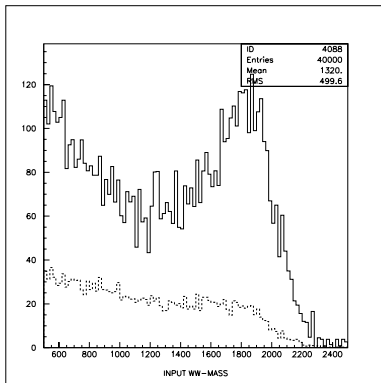


# WW Scattering at CLIC

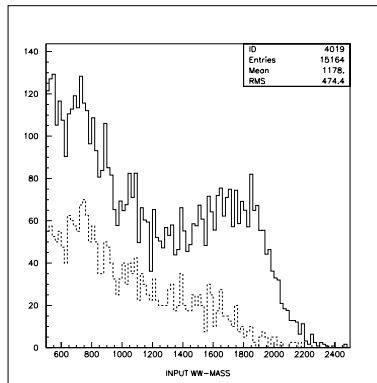
- ◆ Longitudinal  $W$  scattering  $W_L W_L \rightarrow W_L W_L, (t\bar{t})$  <sup>ADR</sup>
- ◆ Analysis of  $WW$  spectrum: excess of events w.r.t. SM and possible resonance structure
- ◆ Study of vector model  $M_{WW} \sim 2$  TeV



EWChL Lagrangian form. ( $a_5 = -0.002, a_4 = 0$ )  
 $1500 \text{ fb}^{-1}$

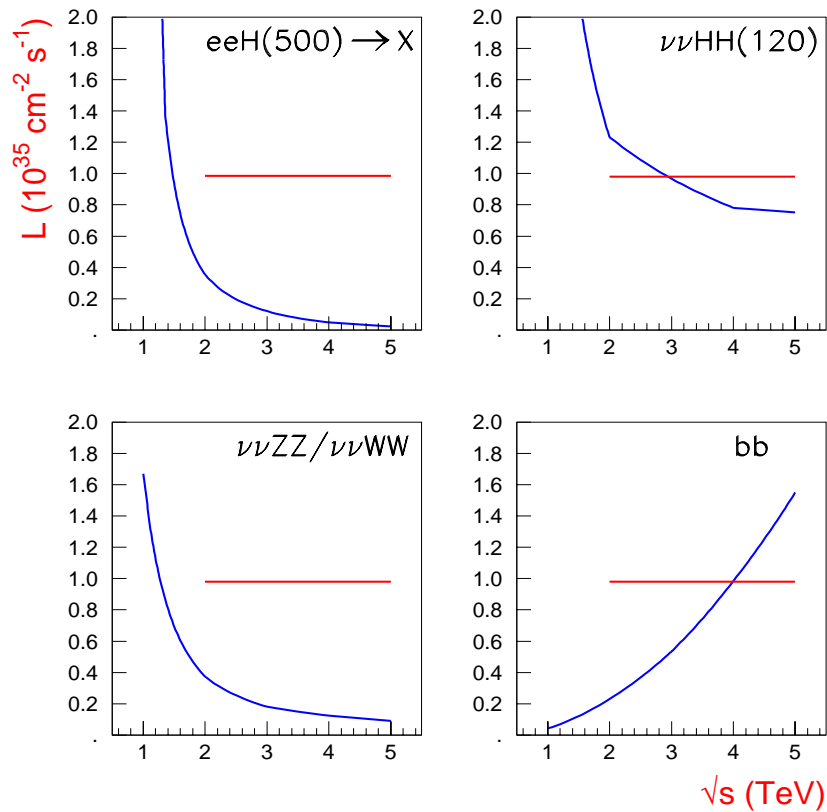


Before detector



After detector

## Luminosity needed for discovery/ precision measurements

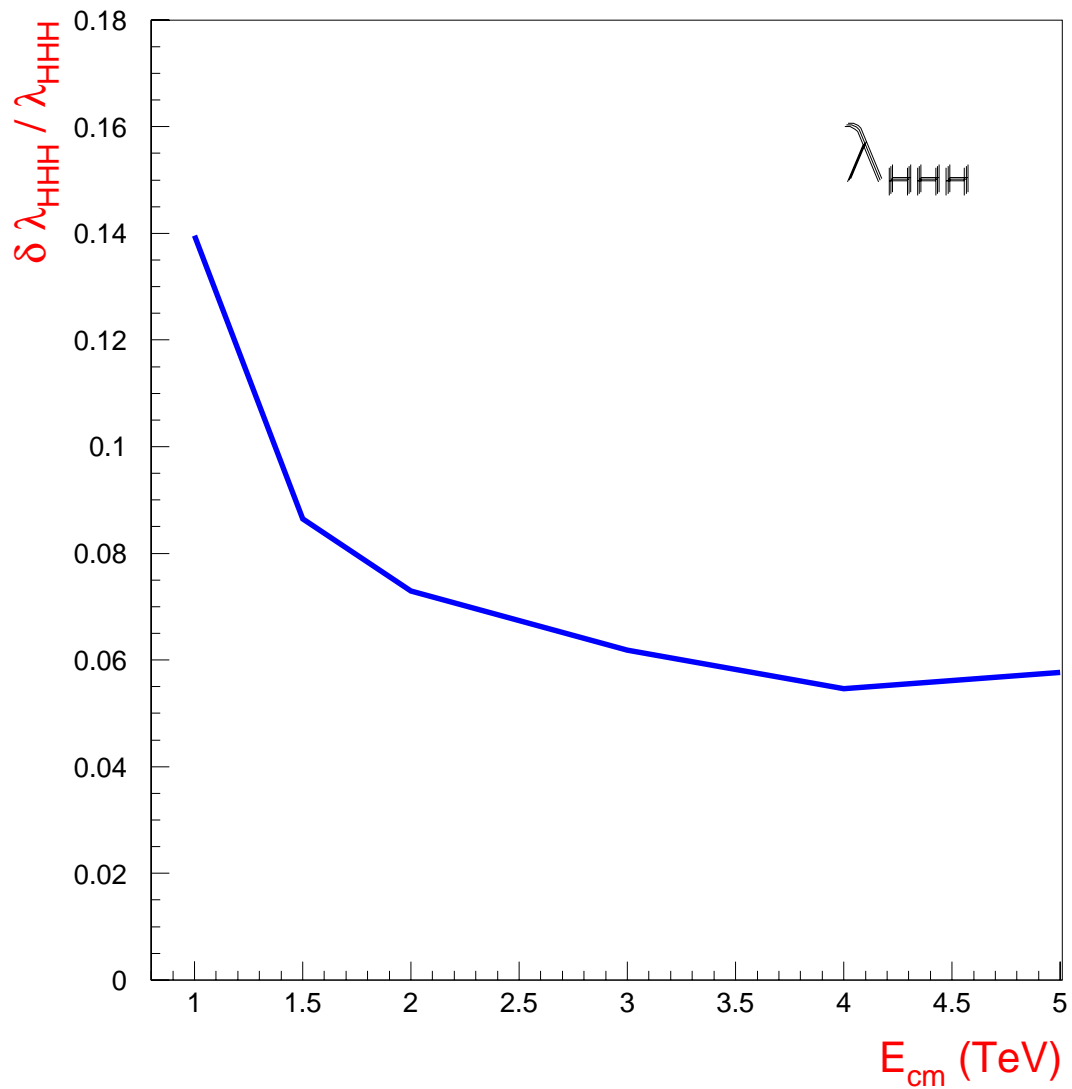


Based on statistical errors, including background

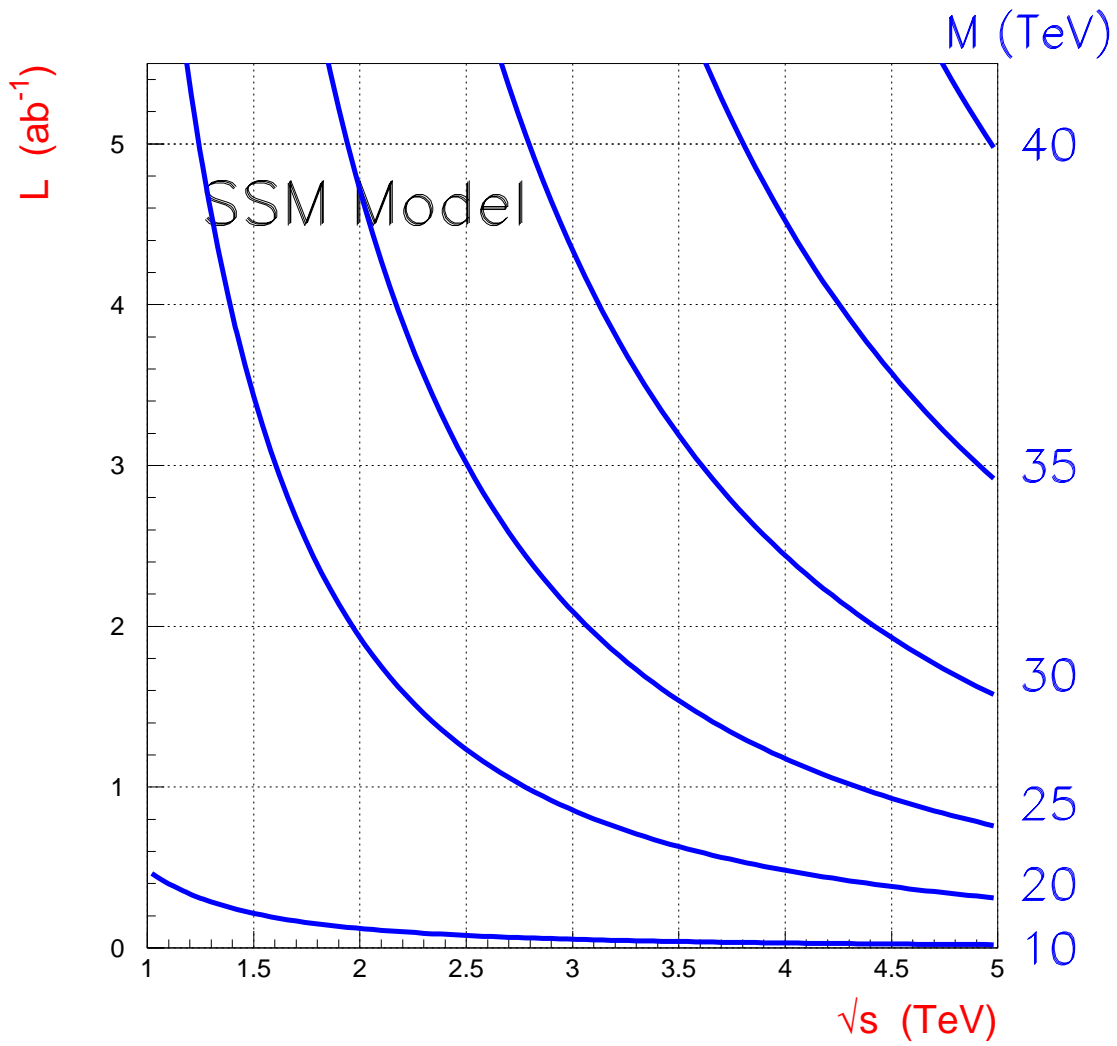
- $e^+e^- \rightarrow H(500\text{GeV, invisible})X$ :  $5\sigma$  discovery
- $e^+e^- \rightarrow \nu\nu HH(120)$ :  $\delta\lambda_{HHH}/\lambda_{HHH} < 15\%$
- $e^+e^- \rightarrow \nu\nu ZZ/\nu\nu WW$ :  $\delta\sigma/\sigma < 1\%$
- $e^+e^- \rightarrow b\bar{b}$ :  $\delta\sigma/\sigma < 1\%$

CLIC Study by A. De Roeck et al.

# Precision on $\lambda_{HHH}$ ( $5000 \text{ fb}^{-1}$ )



Study by M. Battaglia et al.



**95% CL Contours for SSM  $Z'$**   
 (M. Battaglia)

## Sum Rule I

$$\sum_i g_{H_i WW}^2 = g^2 M_W^2$$

### In MSSM

$$g_{hWW} = g M_W \sin(\beta - \alpha)$$

$$g_{HWW} = g M_W \cos(\beta - \alpha)$$

$\cos(\beta - \alpha) \rightarrow 0$  in decoupling limit of  $M_A, M_H \rightarrow \infty$

Can measure  $\Delta \sin^2(\beta - \alpha) / \sin^2(\beta - \alpha) \simeq 3\%$  using  
 $e^+e^- \rightarrow Zh$  and  $e^+e^- \rightarrow \nu\bar{\nu}W^+W^- \rightarrow \nu\bar{\nu}h$  at  
 $\sqrt{s} = 350$  GeV assuming  $M_h = 120$  GeV

(R. Van Kooten)

### Sum Rule I (cont.)

Consider  $W_L W_L \rightarrow Z_L Z_L$  as studied in  $e^+ e^- \rightarrow \nu \bar{\nu} Z Z$

Amplitude  $\mathcal{A}$  for  $W_L W_L \rightarrow Z_L Z_L$  in MSSM :

$$\mathcal{A} = \frac{-s^2(M_h^2 \sin^2(\beta - \alpha) + M_H^2 \cos^2(\beta - \alpha)) + sM_h^2 M_H^2}{v^2(s - M_h^2)(s - M_H^2)}$$

For  $s \gg M_h^2, M_H^2$ :

$$\mathcal{A} = -\frac{1}{v^2}(M_h^2 \sin^2(\beta - \alpha) + M_H^2 \cos^2(\beta - \alpha))$$

For  $s \gg M_h^2, s \ll M_H^2$ :

$$\mathcal{A} = -\frac{M_h^2}{v^2} \left( 1 - \frac{s}{M_H^2} \sin^2(\beta - \alpha) - \frac{s}{M_h^2} \cos^2(\beta - \alpha) \right)$$

For  $s \gg M_h^2, s \ll M_H^2, \cos^2(\beta - \alpha) \rightarrow 0$ :

$$\mathcal{A} = -\frac{M_h^2}{v^2} \left( 1 - \frac{s}{M_H^2} \right)$$

For  $s \gg M_h^2, s \ll M_H^2, \cos^2(\beta - \alpha) \gg M_h^2/s$ :

$$\mathcal{A} = \frac{s}{v^2} \cos^2(\beta - \alpha)$$

## Sum Rule II : Closure of the Neutralino System

$$\begin{aligned}
 \sum_{i,j=1}^4 \mathcal{Z}_{ij} \mathcal{Z}_{ij}^* &= \frac{1}{2} & \sum_{i,j=1}^4 g_{Lij} g_{Lij}^* &= \frac{1}{16c_W^4 s_W^4} \\
 \sum_{i,j=1}^4 \mathcal{Z}_{ij} g_{Lij}^* &= 0 & \sum_{i,j=1}^4 g_{Lij} g_{Rij}^* &= \frac{1}{4c_W^4} \\
 \sum_{i,j=1}^4 \mathcal{Z}_{ij} g_{Rij}^* &= 0 & \sum_{i,j=1}^4 g_{Rij} g_{Rij}^* &= \frac{1}{c_W^4}
 \end{aligned}$$

$$\mathcal{A}(e^+ e^- \rightarrow \tilde{\chi}_i^0 \tilde{\chi}_j^0) = \frac{e^2}{s} Q_{\alpha\beta} [\bar{v}(e^+) \gamma_\mu P_\alpha u(e^-)] [\bar{u}(\tilde{\chi}_i^0) \gamma^\mu P_\beta v(\tilde{\chi}_j^0)]$$

$$\begin{aligned}
 Q_{LL} &= +\frac{D_Z}{s_W^2 c_W^2} (s_W^2 - \frac{1}{2}) \mathcal{Z}_{ij} - D_{uL} g_{Lij} & Q_{RL} &= +\frac{D_Z}{c_W^2} \mathcal{Z}_{ij} + D_{tR} g_{Rij} \\
 Q_{LR} &= -\frac{D_Z}{s_W^2 c_W^2} (s_W^2 - \frac{1}{2}) \mathcal{Z}_{ij}^* + D_{tL} g_{Lij}^* & Q_{RR} &= -\frac{D_Z}{c_W^2} \mathcal{Z}_{ij}^* - D_{uR} g_{Rij}^*
 \end{aligned}$$

(Choi, Kalinowski, Moortgat-Pick, Zerwas, hep-ph/0108117)

## Sum Rule II (cont.)

Neutralino Mass Matrix in  $(\tilde{B}, \tilde{W}^3, \tilde{H}_1^0, \tilde{H}_2^0)$  basis:

$$\mathcal{M} = \begin{pmatrix} M_1 & 0 & -m_Z c_\beta s_W & m_Z s_\beta s_W \\ 0 & M_2 & m_Z c_\beta c_W & -m_Z s_\beta c_W \\ -m_Z c_\beta s_W & m_Z c_\beta c_W & 0 & -\mu \\ m_Z s_\beta s_W & -m_Z s_\beta c_W & -\mu & 0 \end{pmatrix}$$

Do the MSSM neutral gauginos and Higgsinos  $(\tilde{B}, \tilde{W}^3, \tilde{H}_1^0, \tilde{H}_2^0)$  saturate the Neutralino System ?

Chargino Mass Matrix:

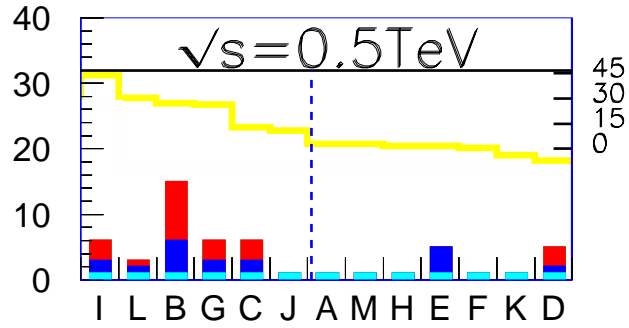
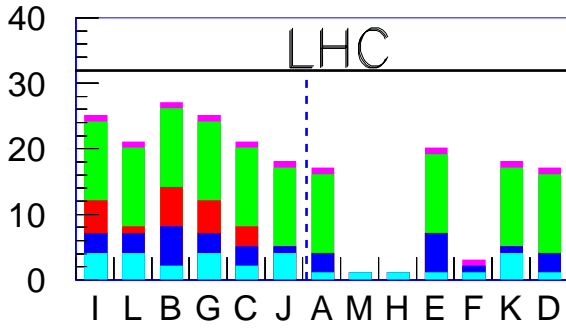
$$\mathcal{M}_C = \begin{pmatrix} M_2 & \sqrt{2} m_W c_\beta \\ \sqrt{2} m_W s_\beta & |\mu| e^{i\Phi_\mu} \end{pmatrix}$$



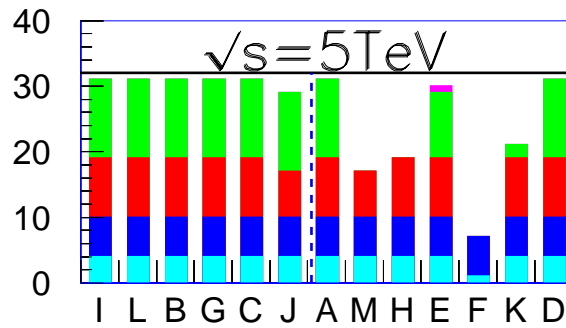
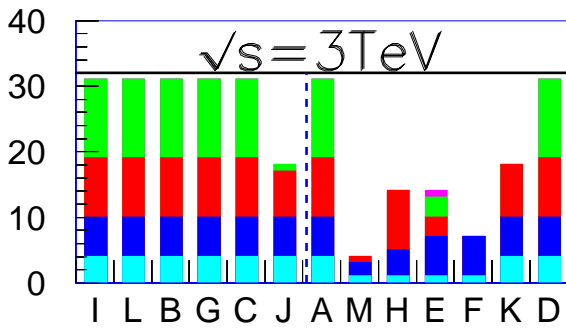
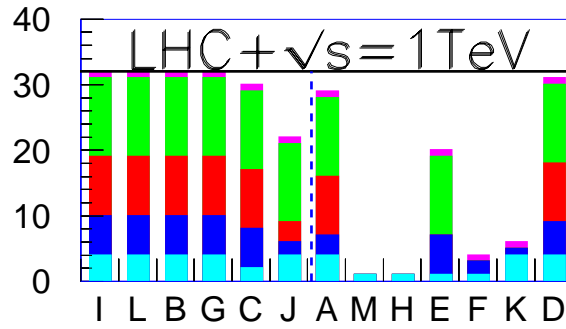
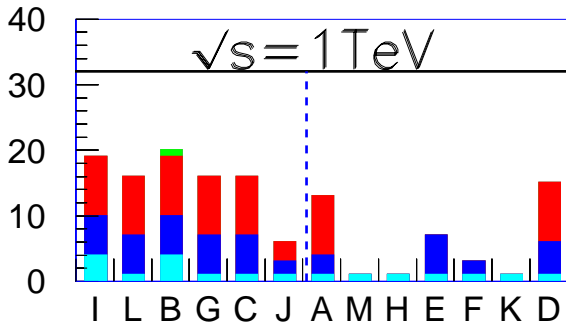
# CMSSM Benchmarks

█ gluino   
 █ squarks   
 █ sleptons   
 █  $\chi^{0,\pm}$    
 █ H

Nb. of Observable Particles



$\delta a_\mu (10^{-10})$   
 45  
 30  
 15  
 0



Post-LEP CMSSM Benchmarks of J. Ellis et al.

Model	A	B	C	D	E	F	G	H	I	J	K	L	M
$m_{1/2}$	600	250	400	525	300	1000	375	1500	350	750	1150	450	1900
$m_0$	140	100	90	125	1500	3450	120	419	180	300	1000	350	1500
$\tan\beta$	5	10	10	10	10	10	20	20	35	35	35	50	50
$\text{sign}(\mu)$	+	+	+	-	+	+	+	+	+	+	-	+	+
$\alpha_s(m_Z)$	120	123	121	121	123	120	122	117	122	119	117	121	116
$m_t$	175	175	175	175	171	171	175	175	175	175	175	175	175
Masses													
$ \mu(m_Z) $	739	332	501	633	239	522	468	1517	437	837	1185	537	1793
$h^0$	114	112	115	115	112	115	116	121	116	120	118	118	123
$H^0$	884	382	577	737	1509	3495	520	1794	449	876	1071	491	1732
$A^0$	883	381	576	736	1509	3495	520	1794	449	876	1071	491	1732
$H^\pm$	887	389	582	741	1511	3496	526	1796	457	880	1075	499	1734
$\chi_1^0$	252	98	164	221	119	434	153	664	143	321	506	188	855
$\chi_2^0$	482	182	310	425	199	546	291	1274	271	617	976	360	1648
$\chi_3^0$	759	345	517	654	255	548	486	1585	462	890	1270	585	2032
$\chi_4^0$	774	364	533	661	318	887	501	1595	476	900	1278	597	2036
$\chi_{1^\pm}$	482	181	310	425	194	537	291	1274	271	617	976	360	1648
$\chi_{2^\pm}$	774	365	533	663	318	888	502	1596	478	901	1279	598	2036
$\tilde{g}$	1299	582	893	1148	697	2108	843	3026	792	1593	2363	994	3768
$e_L, \mu_L$	431	204	290	379	1514	3512	286	1077	302	587	1257	466	1949
$e_R, \mu_R$	271	145	182	239	1505	3471	192	705	228	415	1091	392	1661
$\nu_e, \nu_\mu$	424	188	279	371	1512	3511	275	1074	292	582	1255	459	1947
$\tau_1$	269	137	175	233	1492	3443	166	664	159	334	951	242	1198
$\tau_2$	431	208	292	380	1508	3498	292	1067	313	579	1206	447	1778
$\nu_\tau$	424	187	279	370	1506	3497	271	1062	280	561	1199	417	1772
$u_L, c_L$	1199	547	828	1061	1615	3906	787	2771	752	1486	2360	978	3703
$u_R, c_R$	1148	528	797	1019	1606	3864	757	2637	724	1422	2267	943	3544
$d_L, s_L$	1202	553	832	1064	1617	3906	791	2772	756	1488	2361	981	3704
$d_R, s_R$	1141	527	793	1014	1606	3858	754	2617	721	1413	2254	939	3521
$t_1$	893	392	612	804	1029	2574	582	2117	550	1122	1739	714	2742
$t_2$	1141	571	813	1010	1363	3326	771	2545	728	1363	2017	894	3196
$b_1$	1098	501	759	973	1354	3319	711	2522	656	1316	1960	821	3156
$b_2$	1141	528	792	1009	1594	3832	750	2580	708	1368	2026	887	3216

General Higgs potential with one Higgs doublet:

$$V_{SSB} = \lambda \left\{ \left[ \Phi^\dagger \Phi - \frac{v^2}{2} \right]^2 + \sum_{n \geq 3} \frac{g^{2(n-2)}}{\Lambda^{2(n-2)}} \frac{a_n}{(n-1)^2} \left[ \Phi^\dagger \Phi - \frac{v^2}{2} \right]^n \right\}$$

To leading order in  $1/\Lambda$  the potential is:

$$V_{SSB} = \frac{1}{2} M_H^2 \left\{ H^2 + \frac{g}{M_W} H (\varphi^+ \varphi^- + \frac{\varphi_3^2}{2}) + \frac{g}{2M_W} h_3 H^3 \right. \\ \left. + h_4 \left( \frac{g}{4M_W} \right)^2 H^4 + h'_4 \left( \frac{g}{2M_W} \right)^2 H^2 (\varphi^+ \varphi^- + \frac{\varphi_3^2}{2}) \dots \right\}$$

Non-SM part parameterized by  $\delta h_3$ :

$$h_3 = 1 + a_3 \frac{M_W^2}{\Lambda^2} = 1 + \delta h_3 \\ h_4 = 1 + 6\delta h_3 \\ h'_4 = 1 + 3\delta h_3$$

(F. Boudjema and E. Chopin, hep-ph/9507396)

In MSSM the trilinear coupling of the lightest Higgs:

$$\lambda_{hhh} = -\frac{1}{2v^2} s_{\beta-\alpha} M_h^2 + \frac{c_{\beta-\alpha}^2}{s_\beta c_\beta} \left( s_{\beta-\alpha} \lambda_a + c_{\beta-\alpha} \lambda_b + \frac{(M_A^2 - M_h^2)}{2v^2} (s_{2\beta} s_{\beta-\alpha} + c_{2\beta} c_{\beta-\alpha}) \right)$$

Non-MSSM part parameterized by  $\lambda_a, \lambda_b$

For  $\cos^2(\beta - \alpha) \rightarrow 0$  the trilinear coupling is not sensitive to beyond-MSSM parameters:

$$\lambda_{hhh} = -\frac{1}{2v^2} s_{\beta-\alpha} M_h^2$$

(F. Boudjema and A. Semenov, hep-ph/0201219)

# Summary

- Clear understanding of  $e^+e^-$  physics case beyond  $\sqrt{s} = 0.5$  TeV would help LC technology comparison.
- Analyze specific physics processes at energies of  $\sqrt{s} = 0.8, 1.0, 1.2$  and 1.5 TeV.
- Attempt to identify the energy scale beyond  $\sqrt{s} = 500$  GeV which the next LC should reach through energy expansion.